

Lecture 2.2. **Global Gravity field-modeling.**

T/M Section 6.6

Before 1990 Earth's surface not known precisely: now known from radar altimetry, InSAR and GPS

Before 1990: Problem formulated as Boundary-value problem for elliptic partial differential-equation, boundary S, for volume v.

$$\Delta W = 2\omega^2 = \sum_{i=1}^3 \frac{\partial^2 W}{\partial X_i^2}$$

Solution, if gravity vector component $\left. \frac{\partial W}{\partial n} \right|_S$ vertical to surface, S, is known.

Gravity field modeling.

Anomalous gravity field:

T=W-U,

Observations are linear functionals:

$$\delta g \approx - \left. \frac{\partial T}{\partial r} \right|_P \text{ or } - \left. \frac{\partial T}{\partial h} \right|_P \text{ (better approximation)}$$

$$\Delta g \approx - \left. \frac{\partial T}{\partial r} \right|_P + \frac{T}{\gamma} \left(\frac{\partial \gamma}{\partial h} \right) \approx - \left. \frac{\partial T}{\partial r} \right|_P - \frac{T(P)}{r}$$

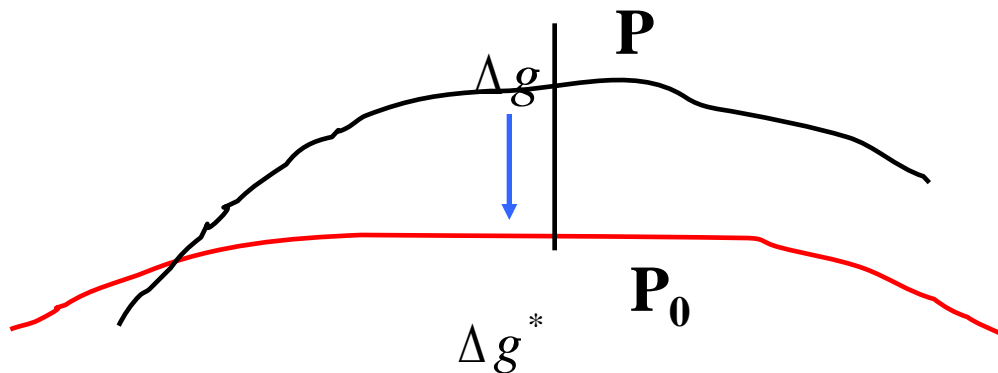
$$\zeta = \frac{T(P)}{\gamma(Q)}$$

Gravity field modeling on Sphere or Ellipsoid.

If Earth may be regarded as spherical or ellipsoidal, then Laplace-equation may be solved using different methods.

Data must be moved from Earth's surface to the Sphere or Ellipsoid and

Masses between surface and Sphere/Ellipsoid must be removed or data corresponding to harmonic function outside sphere/ellipsoid must be determined.



Coefficient-comparison.

Example: polynomium:

$$p(t) = a_0 + a_1 t + a_2 t^2$$

$$p'(t) = a_1 + 2a_2 t$$

If we know that $a_0=0$, then $p(t)$ can be found !

**For anomalous potential are 4 first coefficients = 0,
because we know GM and Earth-center.**

Gravity field modeling, Sphere., p. 273.

$$T(r, \varphi, \lambda) = \sum_{i=2}^{\infty} \frac{GM}{r} \left(\frac{R}{r}\right)^i \sum_{j=-i}^i \bar{P}_{ij}(\sin \varphi) \cdot \begin{Bmatrix} \cos j\lambda \\ \sin|j|\lambda \end{Bmatrix} \Delta \bar{C}_{ij}$$

a put equal to R (mean radius), $r = R + h$, $\bar{\varphi} = \varphi$
then

$$\Delta g(R, \varphi, \lambda) = \sum_{i=2}^{\infty} \frac{GM}{R} \frac{i-1}{R} \sum_{j=-i}^i \bar{P}_{ij}(\sin \varphi) \cdot \begin{Bmatrix} \cos j\lambda \\ \sin|j|\lambda \end{Bmatrix} \Delta \bar{C}_{ij}$$

$$= \sum_{i=2}^{\infty} \frac{GM}{R} \sum_{j=-i}^i \bar{P}_{ij}(\sin \varphi) \cdot \begin{Bmatrix} \cos j\lambda \\ \sin|j|\lambda \end{Bmatrix} H_{ij}$$

$$\Rightarrow \Delta \bar{C}_{ij} = H_{ij} \frac{R}{i-1}$$

Gravity field-modeling.

Spherical harmonic analysis gives H_{ij} and thereby \bar{C}_{ij}

Normalized Legendre-functions used, so that we get orthonormal-system Y_{ij} in space of square-integrable functions $L_2(S)$

$$\frac{1}{4\pi} \iint_S \bar{P}_{ij}(\sin \varphi) \begin{cases} \cos^2 j\lambda \\ \sin^2 j\lambda \end{cases} \cos \varphi d\varphi d\lambda = 1$$

$$Y_{ij}(\varphi, \lambda) = \bar{P}_{ij}(\sin \varphi) \begin{cases} \cos j\lambda \\ \sin |j|\lambda \end{cases} \begin{cases} j \geq 0 \\ j < 0 \end{cases}$$

We use that the coefficients in the development in the space may be determined by calculating the inner product (integral over sphere) of basis-functions with the functions:

$$\frac{GM}{R} \Delta \bar{C}_{ij} = \frac{1}{4\pi} \iint_S T(\varphi', \lambda', R) \cdot Y_{ij}(\varphi', \lambda') dS$$

$$(i-1) \frac{GM}{R^2} \Delta \bar{C}_{ij} = \frac{1}{4\pi} \iint_S \Delta g(\varphi', \lambda', R) \cdot Y_{ij}(\varphi', \lambda') dS$$

Gravity field modeling, Ellipsoid. (EGM2008 !)

$$T(r, \varphi, \lambda) = \sum_{i=2}^{\infty} \frac{GM}{r} \left(\frac{a}{r}\right)^i \sum_{j=-i}^i \bar{P}_{ij}(\sin \varphi) \cdot \begin{Bmatrix} \cos j\lambda \\ \sin |j|\lambda \end{Bmatrix} \Delta \bar{C}_{ij}$$

Then integration over ellipsoid:

$$\Delta \bar{C}_{ij} = \frac{1}{4\pi a \gamma} \sum_{n=0}^{N-1} r_n^e \sum_{k=0}^{s'} \frac{L_{ijk}}{\bar{S}_{n-2k,|j|} (b/E)} \frac{\bar{I}P_{n-2k,|j|}^n}{(i-2k-1)q_{i-2k}^i}$$

$$\sum_{n=0}^{2N-1} \Delta \bar{g}_{nm}^e \begin{Bmatrix} IC \\ IS \end{Bmatrix}_j \begin{matrix} j \geq 0 \\ j < 0 \end{matrix}$$

$$\bar{I}P_{i,|j|}^n = \int_{\delta_n}^{\delta_{n+1}} \bar{P}_{i,|j|}(\cos \delta) \sin \delta d\delta$$

$$\begin{Bmatrix} IC \\ IS \end{Bmatrix}_j^n = \int_{\lambda_m}^{\lambda_{m+1}} \begin{Bmatrix} \cos m\lambda \\ \sin |m|\lambda \end{Bmatrix} d\lambda$$

We use the development for T and integrate over sphere:

$$\begin{aligned}
 T(\varphi, \lambda, r) &= \sum_{i=2}^{\infty} \frac{GM}{r} \left(\frac{R}{r}\right)^i \sum_{j=-i}^i Y_{ij}(\varphi, \lambda) \frac{i-1}{i-1} \left(\frac{R^2}{R^2}\right) \Delta \bar{C}_{ij} \\
 &= \sum_{i=2}^{\infty} \frac{1}{r} \left(\frac{R}{r}\right)^i \sum_{j=-i}^i Y_{ij}(\varphi, \lambda) \frac{R^2}{i-1} \frac{1}{4\pi} \iint \Delta g(\varphi', \lambda', R) Y_{ij}(\varphi', \lambda') dS \\
 &= \frac{R}{4\pi} \iint \left[\Delta g(\varphi', \lambda', R) \sum_{i=2}^{\infty} \frac{1}{i-1} \left(\frac{R}{r}\right)^{i+1} \sum_{j=-i}^i Y_{ij}(\varphi, \lambda) Y_{ij}(\varphi', \lambda') \right] dS
 \end{aligned}$$

Stokes formula, p. 284.

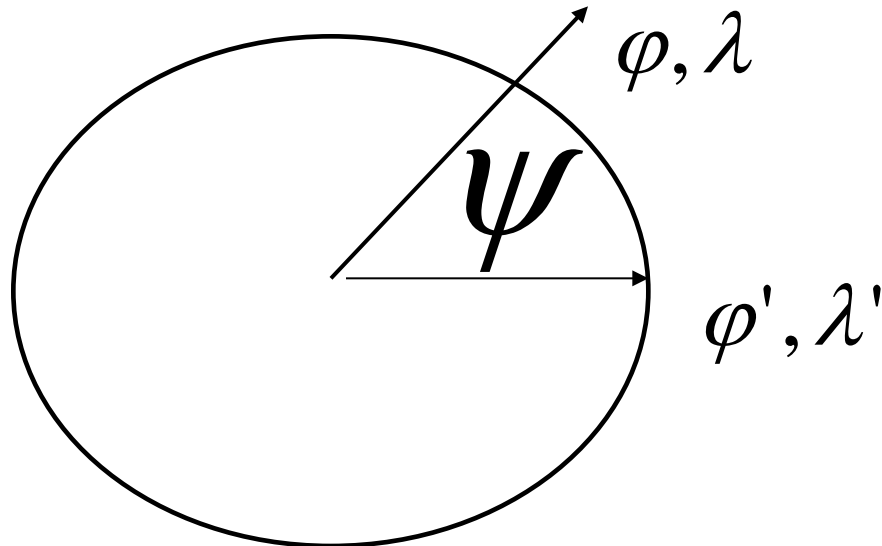
We get:

$$T(\varphi, \lambda, r) = \frac{R}{4\pi} \iint_S \Delta g \cdot \left(\frac{R}{r} \right)^{i+1} \sum_{i=2}^{\infty} \frac{2i+1}{i-1} P_i(\cos \psi) d\sigma$$

Gravity field-modeling.

We use the development in Legendre-polynomials

$$P_i(\cos \psi) = \frac{1}{2i+1} \sum_{j=-i}^i Y_{ij}(\varphi, \lambda) Y_{ij}(\varphi', \lambda')$$



Stokes formula, 1849, Pizetti, 1911. p. 283.

We may obtain closed expression because:

$$\frac{1}{l} = \sum_{i=0}^{\infty} \frac{1}{r} \left(\frac{R}{r} \right)^i P_i(\cos \psi)$$

$$S(\psi, r) = \left(\frac{R}{r} \right)^{i+1} \frac{2i+1}{i-1} P_i(\cos \psi),$$

so on the sphere : $R / r = 1$

$$S(\psi, R) = \frac{1}{\sin(\psi / 2)} + 15 \cos \psi - 6 \sin(\psi / 2) \\ - 3 \cos \psi \cdot \ln(\sin(\psi / 2) + \sin^2(\psi / 2))$$

Stokes formula,

p. 284

We have solved a boundary value problem for an elliptic partial differential equation, but

Note singularity:

$$\frac{1}{\sin(\psi / 2)} \rightarrow \infty$$

when

$$\psi \rightarrow 0$$

Integration must be done with special method.

Height-anomaly/geoid height:

$$\zeta(\varphi, \lambda, r) = \frac{R}{\gamma \cdot 4\pi} \iint_S \Delta g \cdot S(\psi, r) dS$$

Deflections of the vertical, Vening-Meinesz, 1928:

$$\xi(\varphi, \lambda, r) = \frac{1}{\gamma \cdot 4\pi} \iint_S \Delta g \cdot \frac{\partial}{\partial \varphi} S(\psi, r) dS$$

$$\eta(\varphi, \lambda, r) = \frac{1}{\gamma \cdot 4\pi \cos \varphi} \iint_S \Delta g \cdot \frac{\partial}{\partial \lambda} S(\psi, r) dS$$

Stokes formula, implementation. p. 287.

Due to the singularity most of the contribution comes from data with a short distance from the computational-point.

A subtraction of a spherical harmonic model (EGM) to N=360, then only very small area is needed.

$$\zeta(\varphi, \lambda, r) = \zeta_{EGM}(\varphi, \lambda, r) + \frac{R}{\gamma \cdot 4\pi} \iint_{S_0} (\Delta g - \Delta g_{EGM}) \cdot S(\psi, r) dS$$

Note: EGM must be added back later !

Stokes formula, plane Earth. p. 288.

If we want to compute the integral locally, we may regard the Earth as plane.

Integral may be computed using Fourier-transformation (convolution) in the plane, where the spectral relation $(i-1)/r$ is used.

$$\zeta(x, y, R) = \zeta(z, R)_{EGM} + \int \Delta g(t) S(z - t) dt$$

Developed by R.Forsberg and M.Sideris in 1986.

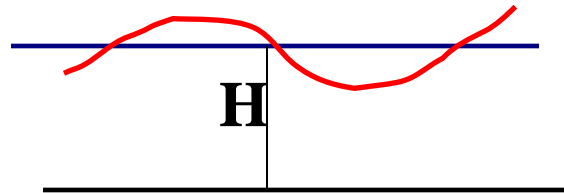
R.Forsberg/CCT: subtract (and add) also local topography

Stokes on the sphere.

Remove masses T_M over altitude 0 (add back later)

Gravity anomaly must then be computed in altitude zero 0
("downward continuation")

Simple method to remove: Bouguer-plate removed in each
point:

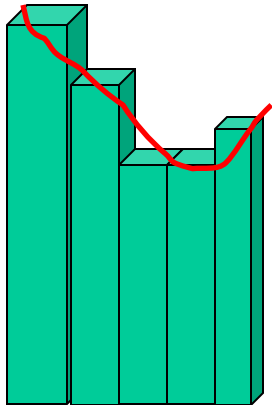


$$2\pi G\rho H \approx 0.11 \text{ mgal} / m$$

Not good because $T - T_M$ not harmonic anymore.

Removal of residual topography, (lecture 3.1) p. 288.

Rectangular prisms better:



May be refined with sloping boundaries – more details in R.Forsberg’s lecture 3.1.

Geodetic Earth-models, see [ICGEM](#).

If only finite number of coefficients used (N) for T and M other data,

then we get traditional least-squares solution with $N+M$ unknowns. Typical $N=240$ or 2196 . $M=300*3+100$ tide-parametre+GM, a , f +datum-shift.

WGS84: NIMA (US DoD),

GEM: Goddard (NASA) Earth Model serie

GRIM: French/German model series

EIGEN: German/French/US, latest EIGEN6_C2

ITG-Grace: Uni. Bonn.

GOCE: GO_CONS_GCF_2_DIR/TIM_R3/4.

Only gravity: OSU91, EGM96, EGM2008.